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INTERNATIONAL ASTRONOMICAL UNION

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COMET P/2008 L1 (LARSEN)

R. S. McMillan, Lunar and Planetary Laboratory, University of Arizona, reports that J. V. Scotti has recovered comet P/1997 V1 (cf. *IAUC* 6767) with the Spacewatch 1.8-m $f/2.7$ reflector at Kitt Peak (recovery observation tabulated below). On June 9.4 UT, Scotti described the coma as of diameter $9''$, and there was a faint tail extending $0'.15$ in p.a. 243° ; on June 10.4, he gave the coma diameter as $10''$, the tail extending $0'.11$ in p.a. 252° . The indicated correction to the prediction on *MPC* 54170 (ephemeris on *MPC* 62064) is $\Delta T = -2.0$ days.

2008 UT	α_{2000}	δ_{2000}	Mag.
June 9.44185	$0^{\text{h}}25^{\text{m}}34.60^{\text{s}}$	$+13^{\circ}29'14.0''$	20.2

The available astrometry, the following orbital elements (along with revised elements for 1997), and an ephemeris appear on *MPEC* 2008-L45.

Epoch = 2008 Sept. 11.0 TT			
$T = 2008$ Aug. 25.1744 TT	$\omega = 133.7286$	} 2000.0	
$e = 0.333226$	$\Omega = 234.8163$		
$q = 3.271999$ AU	$i = 12.1214$		
$a = 4.907208$ AU	$n^\circ = 0.0906677$	$P = 10.87$ years	

V5579 SAGITTARII

R. J. Rudy, D. K. Lynch, R. W. Russell, K. Crawford, and B. Kanehiro, Aerospace Corporation; C. E. Woodward, University of Minnesota; M. Sitko, University of Cincinnati and Space Science Institute; and M. Skinner, Boeing LTS, report on spectroscopic observations (range 0.8 – 13.5 μm) of the nova V5579 Sgr from May 22 UT. Data from 0.8 – 5.2 μm were obtained using the SPEX instrument at the Infrared Telescope Facility, while measurements from 3 – 13.5 μm were acquired with the Advanced Electro-Optical Telescope using the Aerospace Corporation's Broadband Array Spectrograph System. The nova has changed dramatically since the early observations from May 9 (*IAUC* 8948). The dust emission has increased and now dominates the continuum. The equivalent widths of the lines have weakened substantially, and the line profiles now show a pronounced absorption at their centers. The excitation of the emission lines remains low, although the He I line at 1.083 μm is now present. The dust has cooled significantly and now has a temperature of 1080 K.